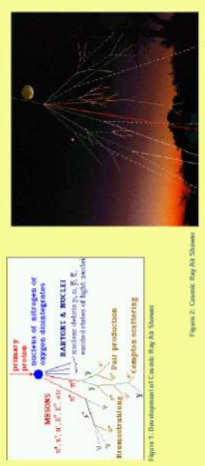


# Cosmic Ray Muon Detection

## Measurements of Cosmic Ray Muon Flux and Muon Lifetime Using Scintillation Detectors

G. Karagiorgi, J. Slanker, Dr. M. Hohlmann, Florida Institute of Technology

Muons are created in the upper atmosphere of the Earth by cosmic rays, which are mostly protons that come from outer space. When a high-energy proton hits one of the nuclei of the atmospheric gas molecules, it creates an "air shower," in an air shower, different subatomic particles are created, including pions, which then decay into muons. These highly energetic muons interact only weakly with the air and have a mean lifetime long enough for some to arrive at the Earth's surface, where they can be detected.



### Project Summary

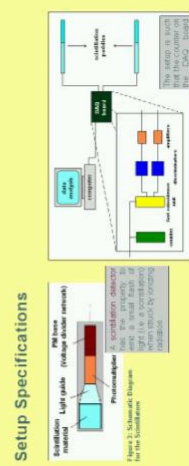
Using a setup of two scintillation detectors:

- Flux
- Count rate
- Energy variation of muons originating from cosmic ray air showers are investigated.

The factors considered are:

- Amount of material muons travel through
- Zenith angle
- North-South (NS) vs. East-West (EW) orientation
- Setup configuration of the detection system (overlap area and separation distance)

Additionally, the method of detection allowed for verification of the theoretical value for the muon lifetime.



### Background Concepts

**Flux**  
Muons reach the surface of the Earth with typically constant flux  $F_{\mu}$ :

$$F_{\mu} = \frac{\text{(count rate)} \cdot d^2}{\text{(area of top panel)} \cdot \text{(area of bottom panel)}}$$

(horizontal detector)  
 $F_{\mu} = 0.04 \text{ cm}^{-2} \text{ min}^{-1}$  (measured) (PDG theoretical value)  
 Our experimental value:  $3.9 \text{ cm}^{-2} \text{ min}^{-1}$  (8% efficiency)

The flux varies with zenith angle  $\theta$  as:

$$F_{\mu} = \cos^2 \theta$$

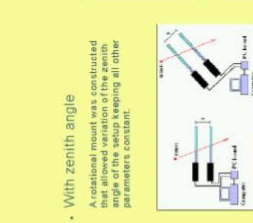

### Results

#### Investigation of Flux Variation

- With material above detectors

Data were collected on the 7 different floors\* of Crawford Building, on the Florida Tech Campus

\* 7 floors (including the 7th floor) were used for this data set (the 7th floor)



#### With zenith angle

An additional mount was constructed that allowed variation of the zenith angle of the setup keeping all other parameters constant.



#### With NS vs. EW orientation

Data were collected for a 180-degree rotation of the setup for both NS and EW orientation.



#### With overlap area

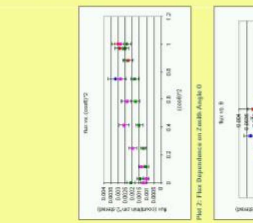
The measurement was conducted to confirm the independence of flux with detector non-uniformities in the active area of the scintillation paddles.



#### Investigation of Count Rate Variation

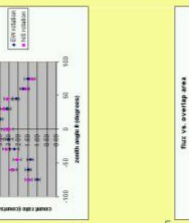
- With separation distance  $d$  between the two paddles

The paddles were placed in a rectangular arrangement. The active area (horizontal) was kept constant ( $d^2$ ), and the separation distance  $d$  was altered in multiples of 1.



#### Investigation of Energy Variation

Using the Counter DCO v1 board, the energy (displaying muon events on the computer). These events are called "doubles."

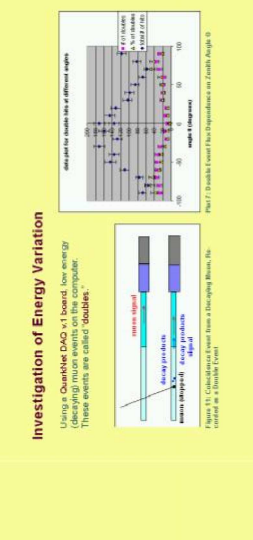
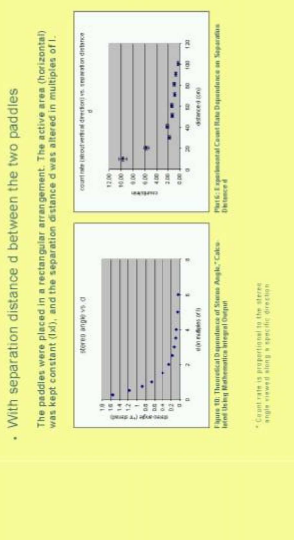
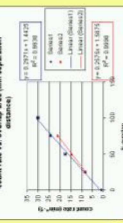


#### Muon Lifetime Measurement

The muon lifetime was calculated using the Counter DCO v1 board data for double events. The data was plotted and the data were fitted to an exponential curve of the form:

$$N(t) = N_0 e^{-t/\tau}$$

where  $\tau$  = muon lifetime

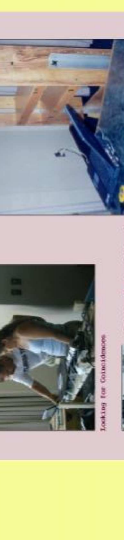


#### Muon Lifetime Measurement

The muon lifetime was calculated using the Counter DCO v1 board data for double events. The data was plotted and the data were fitted to an exponential curve of the form:

$$N(t) = N_0 e^{-t/\tau}$$

where  $\tau$  = muon lifetime



Presentations will be available for discussion from 8:30am to 7:30pm